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CELESTITE FROM PERMIAN DEPOSITS OF THE SOUTHERN TIMAN

D. P. Serdyucherko Presented by Acad D. S. Belyankinyy Received 20 June 1947

During a 1941 survey of samples from bore hole No 1 at Bol shipe Porogi in Vysskiy Rayon (Komi ASSR), among the rooks of the lower Permian anhydro-colomite stratum at a death of about 500 meters, I discovered celestite in the form of massive, coarse (3hy 4h/3 mantimatems) crystal aggregates inclosed in massive fine-grained light-gray dolomite.

The celestite is gray-blue or sky-blue in color. The mineral is transparent or senitransparent. The microscope indicates perfect cleavage along face 010 and less perfect cleavage along face 001. Since cleavage occurs at right angles, celestite can be split into elongated rectangular little blocks. Frismatic cleavage along face 110 is perfect and forms an orthor remain; system with the acute angle between cleavage cracks 110 and 110 equal-ling 75 degrees. This measurement in Fedorovakiy's table does not differ greatly from the goniometric measurement by which this angle (according to Dama) equals 75 degrees, 50 minutes for celestite.

Hg equals 1.630 $\stackrel{\downarrow}{=}$ 0.002; Hm equals 1.624 $\stackrel{\downarrow}{=}$ 0.002; Hp equals 1.621 $\stackrel{\downarrow}{=}$ 0.002; Hg - Hp equals 0.009; 2V equals 49.5 degrees; $\rho \ll \nu$. The plane of the optic axes is 010.

Extinction of the mineral with respect to the cleavage planes along OlO and OOl is parallel, but with respect to prismatic cleavage along ll¢ it is symmetrical. Specific gravity, determined by pycnometer at 20 degrees, equals 4.01.

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Chemical Composition of Mineral

	Persent	Molscular Part
8102	0.64	
MgO	0.30	0.008
CaO	1.12	0.020
SrO	53.76	0,519
BaO	0.42	0.003
so ₃	43.95 100.19	0.548

Taking into consideration the sizes of the ionic radii (radiusov) Mg (0.78 Å), Ca (1.06 Å), Sr (1.27 Å), and Ba (1.43 Å) and also microscopic inclusions of dolomite in celestite crystals, we combine the total determined by analysis) Mg with the corresponding amount of Ca in a molecule of dolomite.

Finally we obtain the composition of celestite: CaSO₄-2.25, SrSO₄-97.19, BaSO₄-0.56, total 100.00 percent. This corresponds with the formula (Ca, Sr, Ba) SO₄.

Paragenesis of celestite is established by its common occurrence and intimate association with dolcmite, anhydrite, and gypsum.

Dolomite is massive light-gray rock shown under the microscope to consist of very fine (0.05-0.005) rhombohedral, or point, grains which impart pelitomorphic structure to the rock. In cross sections of celestite, especially those perpendicular to the prism, inclusions of dolomite are clearly visible. These are sometimes dispersed as single fine grains, and sometimes occur as fine-grained relitomorphic spotty masses.

Anhydrite interstratifies with dolomite. It occurs in the form of strata, streaks, or lenticular bodies. The color of anhydrite is azure or azure-blue. The mineral has a glassy and vitreous transparen y. Sometimes it has a saccharoidal or fine-grained structure. Under the microscope are seen very elongated prismatic fibers, which are curved, and either radiating or crisscrossing threads and strands with cleavage planes transverse to their lengths. Occasionally, masses of fine dolomite grains occur along the fibers. The microstructure of anhydrite rock is fibrov (nematoblastic).

Cleavage of the mineral is perfect along facts 010 and 001, less perfect along 100; extinction is parallel; 2V equals \$1.5 degrees; 2V lies on face 010; Mg equals 1.615; Mg equals 1.572; Mg-Fp equals 0.0Mg.

Specific gravity, determined by pyonometer as 20 degrees centigrade, varies within the limits 2.88-3.13 for different samples.

Gypsum occurs in the form of the comparatively coarse idioblastic monoelimic crystals typical of this mineral, which are sometimes in the twimning position of the "dovetail" (lastochkin khvost) type. Gypsum monocrystals usually occur in small number in the immediate vicinity of dolomite, but develop partly at the expense of anhydrite fibers which they replace: 2V equals + 56 degrees; Ing - Inp equals 0.009; Ng<1.537; 2V lies on face 010.

The genesis of celestite in the lower Permian deposits of Vymskiy Rayon is clearly connected with its precipitation by chemical means from highly con-

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centrated lagoon or residual salt-lake water, a fact which its entire mineral paragenesis indicates. In other similar deposits it is also common to find celestite associated not only with dolomite, anhydrite, and gypsum, but also with fluorite, halite, and sylvite.

Apparently, finely dispersed celestite crystals in the process of diagenesis and collective recrystallization originally formed comparatively coarse crystals and aggregates in pelitomorphic dolomite deposition.

The Vymskiy Rayon celestite deposit is connected with the post-Hercynian period of deposition and gravitates toward the extensive Permian celestite range which, according to L. V. Pustovalov, "is detected from the Horthern region (Pinega River and Severnaya Dvina River basins) through the Tatar Republic up to the Kuybyshev region (middle Volga, vicinity of city of Kuybyshev, etc.) and enters into the Pre-Ural region (Bashirt ASSR, Chialev Oblast) and even into the Urals (eastern Ural slope, vicinity of city of Emmesk); in all these places celestite is associated with gypeiferous rocks of the Artinekian, Kungurekian, or upper Permian age."

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